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Final Technical Report
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Analysis of Magnetic Field Data from
the Pioneer Venus Orbiter

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Introduction

The subject grant NAG 2-501 supported the analysis of the magnetic field data from the Pioneer Venus orbiter for the period 10/1/87 to 9/30/94. During that period 188 papers were contributed to scientific meetings that either analyzed the magnetometer data or used the magnetometer data as part of the analysis of a scientific problem. Similarly, 107 papers were published in research journals and books. Members of the team were asked to give 15 invited review papers on the results of the mission and four theses were prepared using these data. These papers and theses are listed in the tables at the back of this report. Table 1 lists the papers presented at meetings analyzing the magnetometer data. Table 2 lists the corresponding papers in journals and books. Table 3 lists the papers presented at meetings in which the magnetometer data was used in a correlative role. Table 4 lists the corresponding papers in journals and books. Table 5 lists the papers invited for presentation at scientific meetings. Table 6 lists the theses analyzing these data.

A brief summary of the work contained in the published papers of Table 2 can serve as a summary of the work accomplished under this grant. We will begin at lowest altitudes and work outwards. The magnetic moment of Venus was discussed in papers (2.1) and (2.55). We found Venus to be essentially devoid of any intrinsic magnetic field. In contrast we found much evidence for the presence of lightning in the Venus atmosphere (2.26, 2.28, 2.32, 2.36). We mapped the altitude distribution (2.8), found geographic cluster (2.13) most probably associated with local time ordering (2.21, 2.23). Some of the waves were clearly electromagnetic propagating below the electron gyro frequency but others were not (2.13, 2.25). For those waves that were electromagnetic we measured the Poynting flux (2.17). We examined the entry data at low altitudes (2.61) and the magnetization of the ionosphere and the atmospheric region below (2.3, 2.6, 2.38, 2.45, 2.49, 2.54, 2.62). We looked at flux ropes in the ionosphere (2.29) and postulated a new means to create them (2.31). On the nightside we studied ionospheric holes (2.50), ionospheric clouds (2.41) and tail rays (2.41). We studied the subsolar ionopause (2.2) and the magnetic barrier (2.42). We examined the altitude asymmetry of the ionopause (2.7), the properties of the magnetosheath (2.52) and the location of the bow shock (2.12, 2.14, 2.24, 2.37, 2.47, 2.48, 2.58) as well as the properties of upstream ions (2.18) and upstream waves (2.35, 2.44, 2.56, 2.57, 2.64). We examined the solar wind interaction overall (2.30, 2.40) and the properties of the magnetotail in particular (2.15, 2.33, 2.39, 2.43, 2.47, 2.59, 2.60). We also used the Pioneer Venus data to learn more about the processes occurring in the solar wind (2.4, 2.16, 2.19, 2.27, 2.51, 2.53, 2.63).

No patents or inventions resulted from this grant. All data obtained by the magnetic field investigation whether raw or processed have been documented and submitted to the Planetary Data System for archiving.

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Table 1

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